

Compositional Diversity in the South Pole-Aitken Basin (SPA) as Viewed by the Moon Mineralogy Mapper (M³)

Noah Perto, Carlé Pieters, Rachel Klima and the M³ Science Team

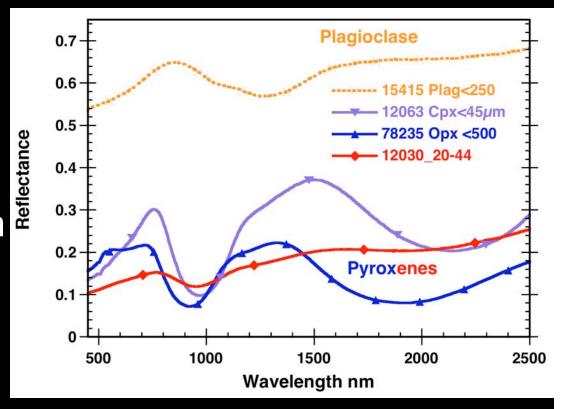






M³ Instrument Overview

- M³ was designed to characterize the diagnostic absorption features of common lunar minerals
- 140 m/pixel at 100 km orbit (but....)
- 86 spectral channels from 430 to 3000 nm



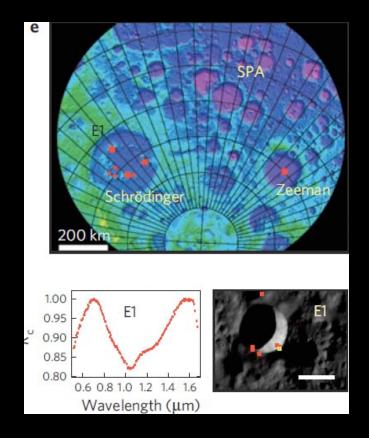


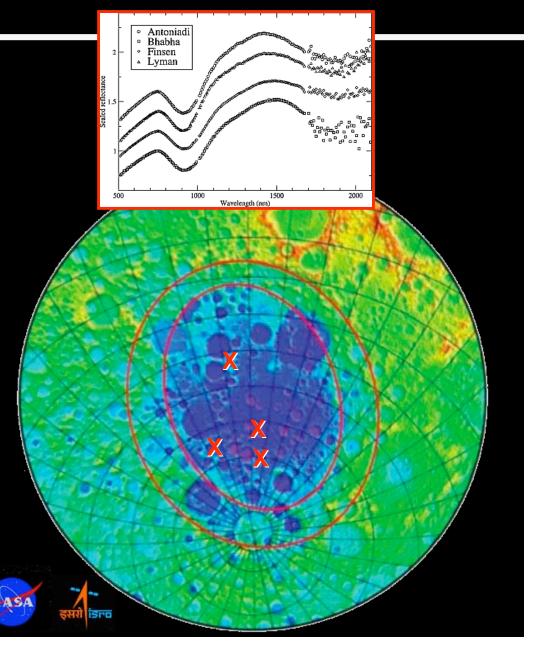




Kaguya Spectral Profiler Results

The central peaks of several large craters exhibit prominent features due to (crystalline) Mg-rich pyroxene. (Nakamura et al., 2009) Olivine in Schrödinger and Zeeman. (Yamamoto et al., 2010)







What did we get? OP2A @ 100km















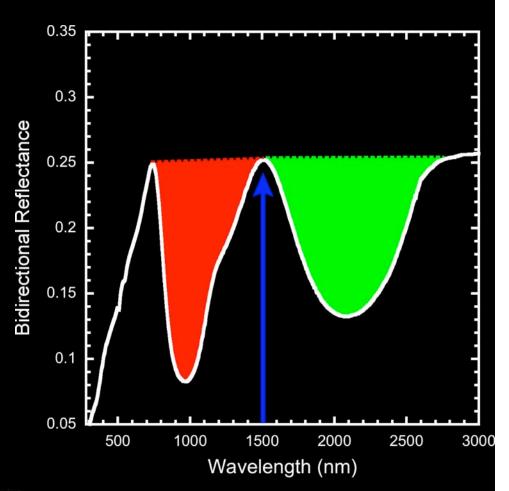






Strength of M³

- Single bands are great, but the strength of M³ comes from leveraging the spectral resolution
- Useful color composite leverages the integrated 1 and 2 micron band depths, and the albedo at 1580 microns

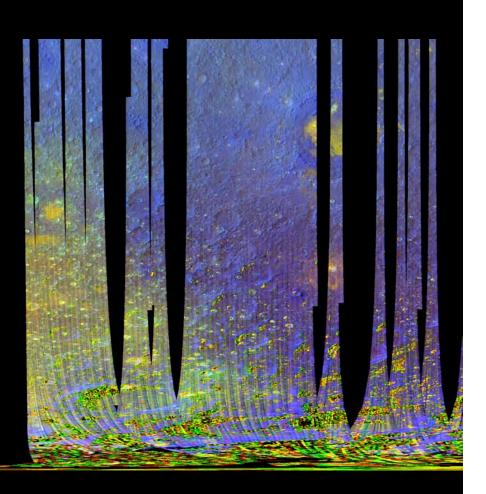








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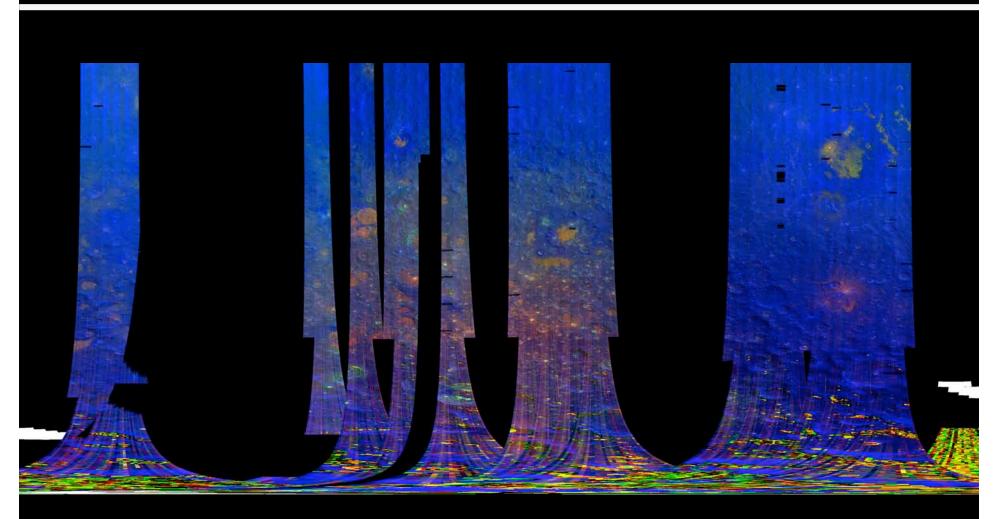






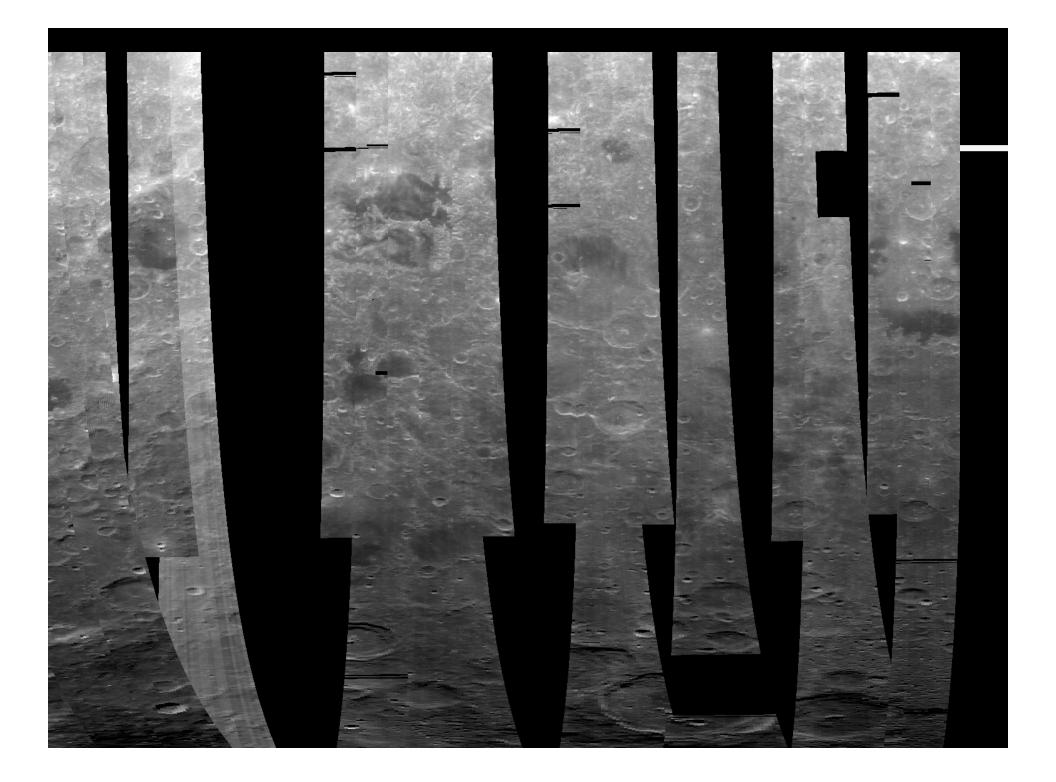


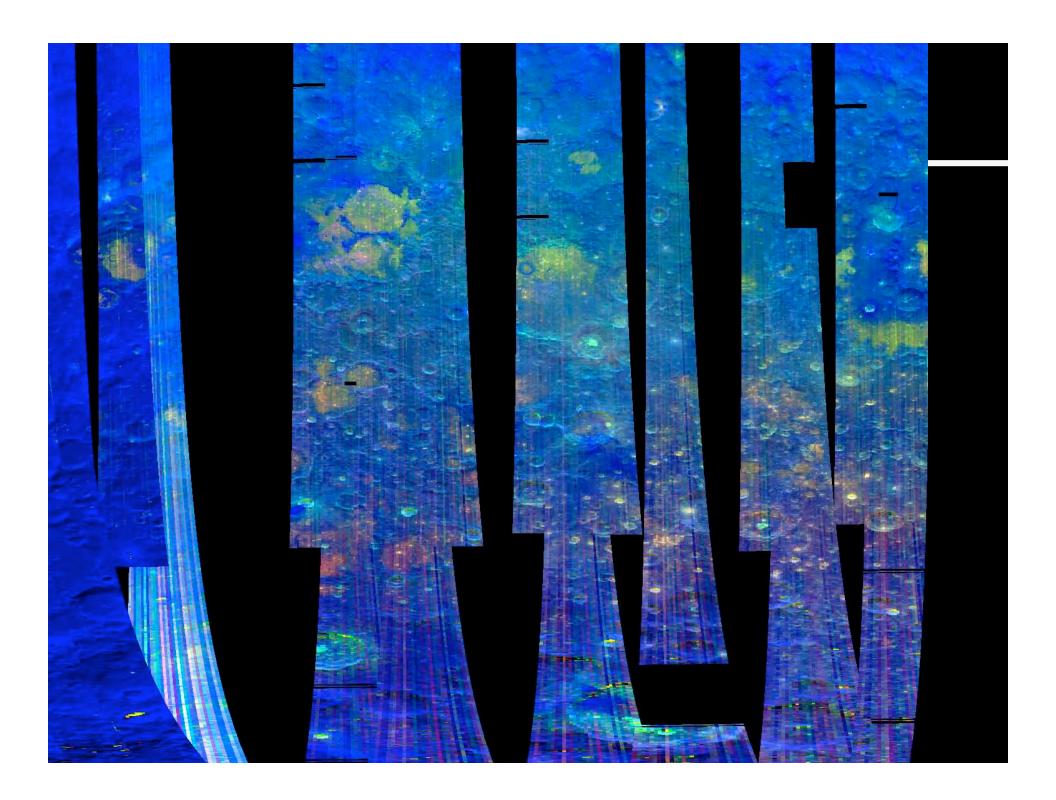


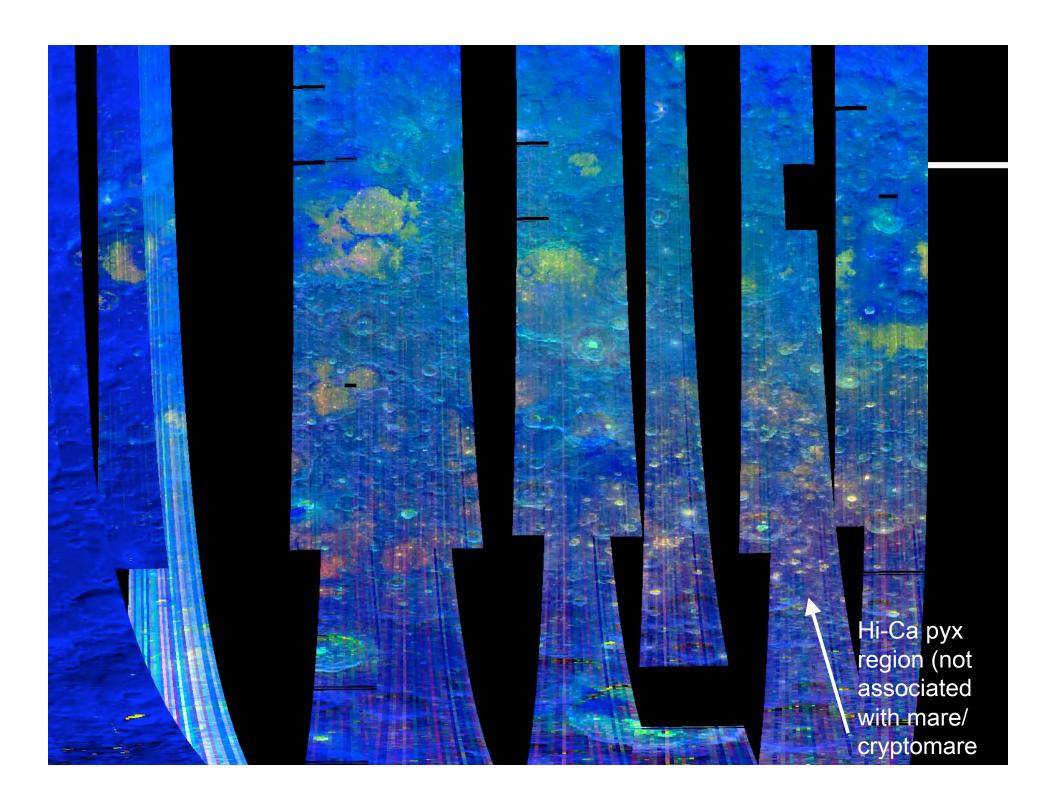


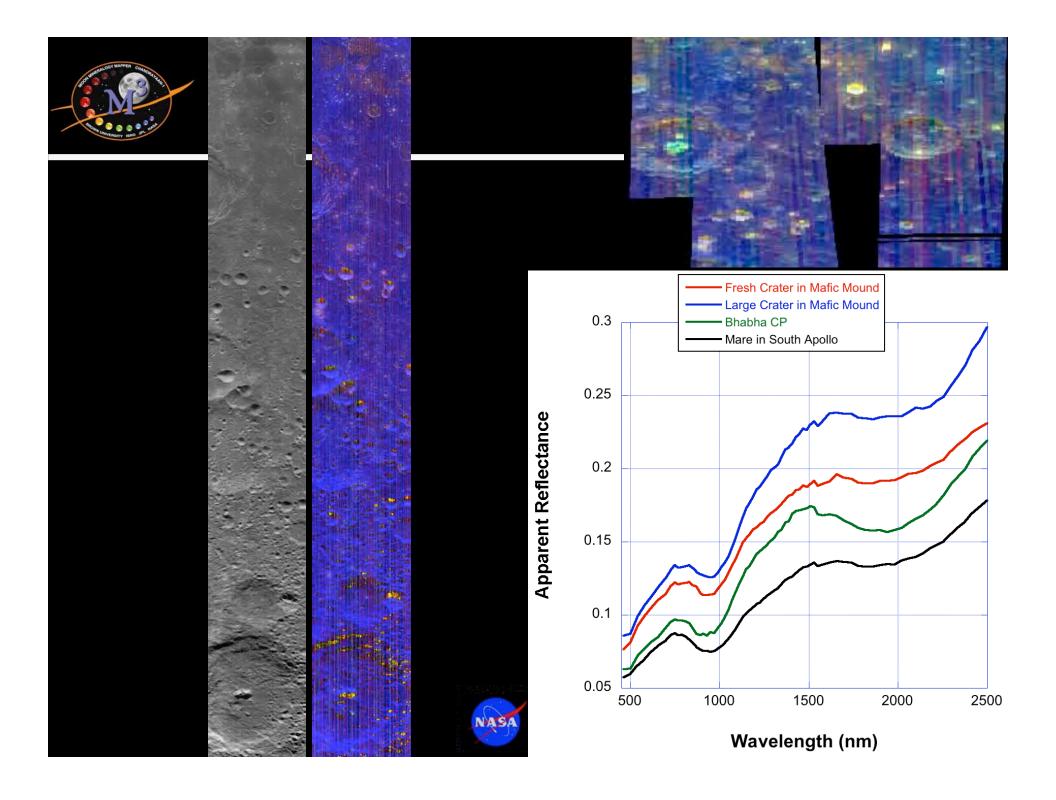






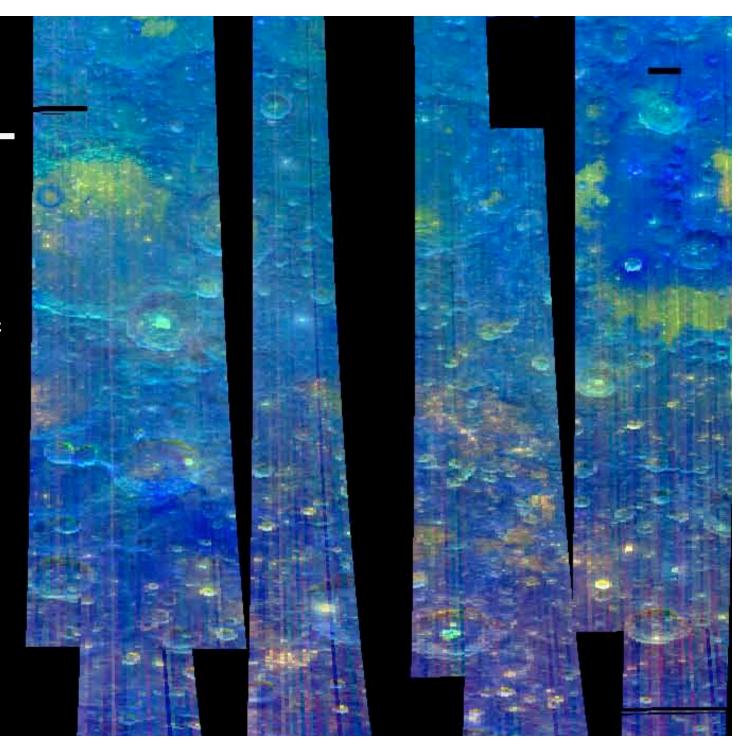








- Noritic material in Apollo is compositionally distinct (higher Mg) than noritic material outside Apollo
- High-Ca pyx. in middle of basin, over noritic material?





Current SPA Story...

- The highly noritic rock type seen in the central peaks of large craters represents excavated lower crustal material (below the impact melt) or a is derived from SPA impact melt (e.g., Nakamura et al., 2009).
- The rest of the mafic-rich SPA basin interior largely represents products of the SPA impact melt that have been reworked with time.
 - The moderately **noritic** SPA interior represents a mixed impact melt breccia of the crustal column produced during the impact event.
 - The gabbroic zone across the center of the basin a distinct portion of the impact melt and has not been homogenized with the rest. Overlies noritic material (Bhabha Central Peak)

The low-Fe bearing materials in SPA represent....











Low-FeO Material

